

Principle of virtual action in continuum mechanics

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Abstract We present the principle of virtual action as a foundation of continuum mechanics. Used mainly in relativity, the method has a useful application in classical mechanics and places the notion of action as the basic concept of dynamics. The principle is an extension of virtual work to space-time. It extends the efforts made by d'Alembert and Lagrange. Unlike the classical case of equilibrium, the principle of virtual action becomes a postulate for the formulation of models in dynamics. It allows to use a minimal set of clear conjectures and is extended to the case of media with dissipation; it can be used for more complex systems.

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1 Introduction

For continuous media, the principle of virtual work is convenient to obtain the equations of equilibrium [1, 2, 3, 4, 5]. Forces and more generally stresses are dual of displacements and the principle yields equilibrium equations. It can also be used to obtain the equations of conservative motions: it is enough to add inertial forces to the different efforts.

However, since motions are associated with space-time quantities, it is natural to use the notion of action rather than the notion of energy [6, 7, 8]. Indeed, all quantities involved in dynamics such as momentum and inertial forces are derived from the action. A principle of virtual action is more efficient than the principle of virtual work because it directly gives not only the equation of momentum but also the equation of energy. The thermodynamics being introduced with the greatest simplicity, we naturally obtain the equation of entropy. This way of thinking, not usual in classical mechanics, is the basis of

general relativity which uses same concepts as classical mechanics but differs in the choice of dynamic behavior laws [9].

The motion of a continuous medium is governed by an application φ_t from $\mathcal{D}_0$ into $\mathcal{D}_t$, where $\mathcal{D}_0$ is an open set of a three dimensional domain and $\mathcal{D}_t$ is at time t an open set of the three-dimension physical space $\mathcal{E}$ occupied by the medium [1]; $\mathbf{X} = [X^1, X^2, X^3]^T$ denotes the reference position in Lagrange variables, $\mathbf{x} = [x^1, x^2, x^3]^T$ denotes the particle position in Euler variables at time t , where T denotes the transposition; we write

$$\mathbf{x} = \varphi_t(\mathbf{X})$$

We can also write the motion in the form

$$\begin{cases} t = t(\lambda, \mathbf{X}) \\ \mathbf{x} = \mathbf{x}(\lambda, \mathbf{X}) \end{cases} \quad \text{or} \quad \mathbf{z} = \mathcal{F}(\mathbf{Z}) \quad (1.1)$$

where λ is a scalar parameter and

$$\mathbf{z} = \begin{bmatrix} t \\ \mathbf{x} \end{bmatrix}, \quad \mathbf{Z} = \begin{bmatrix} \lambda \\ \mathbf{X} \end{bmatrix},$$

with $z^0 = t$, $z^i = x^i$, $Z^0 = \lambda$, $Z^i = X^i$, $i \in \{1, 2, 3\}$.

Between times t_1 and t_2 , the current domain $\mathcal{D}_t$ of element $[t, \mathbf{x}]^T$ generates a manifold $\mathcal{W}$ of the 4-dimension affine space-time occupied by the medium and $\mathcal{W}_0 = \mathbb{R} \times \mathcal{D}_0$ becomes the parameter space. The motion $\mathcal{F}$ of the continuous medium, corresponding to a diffeomorphism on manifold with edge, is defined on an open set so that $\mathcal{W}$ and $\mathcal{W}_0$ have compact closures and piece-wise differentiable boundaries [4, 7]. We note the similarity with the motion defined by an application between two 4-dimensional spaces and the statics position defined by an application between two 3-dimensional spaces.

2 Principle of virtual action

2.1 Virtual motions and virtual displacement of a continuous medium

Definition 1 Consider a differentiable application $\mathcal{G}$ such that

$$(\mathbf{Z}, \varepsilon) \in \mathcal{W}_0 \times \mathcal{O} \longrightarrow \mathbf{z} = \mathcal{G}(\mathbf{Z}, \varepsilon) \in \mathcal{W} \quad \text{with} \quad \mathcal{G}(\mathbf{Z}, 0) = \mathcal{F}(\mathbf{Z})$$

where $\mathcal{O}$ is an open set of $\mathbb{R}$ containing 0. We call $\mathcal{G}$ a one-parameter family of *virtual motions*.

The associated *virtual displacement* is the vector field

$$\mathbf{z} \in \mathcal{W} \longrightarrow \boldsymbol{\zeta} \quad \text{where} \quad \boldsymbol{\zeta} = \frac{\partial \mathcal{G}(\mathbf{Z}, 0)}{\partial \varepsilon},$$

which belongs to the vector space, tangent to $\mathcal{W}$ at point $\mathbf{z} = \mathcal{F}(\mathbf{Z})$. In fact we can also write $\boldsymbol{\zeta}$ as a differential of $\mathcal{G}(\mathbf{Z}, \varepsilon)$ when $\varepsilon = 0$

$$\boldsymbol{\zeta} = d\mathbf{z} \quad \text{with} \quad d\mathbf{Z} = \mathbf{0} \quad \text{and} \quad d\varepsilon = 1$$

A linear functional of the virtual displacement is called virtual action and is denoted by δa . To describe the most classical model, we choose the expression

$$\delta a = \int_{\mathcal{W}} \left[S_i \zeta^i + T_i^j \zeta_{,j}^i \right] dw$$

where dw is the *volume element* of $\mathcal{W}$. In matrix notation, we obtain

$$\delta a = \int_{\mathcal{W}} \left[\mathbf{S}^T \boldsymbol{\zeta} + \text{Tr} \left(\mathbf{T} \frac{\partial \boldsymbol{\zeta}}{\partial \mathbf{z}} \right) \right] dw \quad (2.2)$$

which introduces a field of covectors $\mathbf{S}^T$ (or forms) and a field of matrices $\mathbf{T}$ which are separated into time and space parts

$$\mathbf{S}^T = [-\mathcal{P}, \mathbf{F}^T] \quad \text{and} \quad \mathbf{T} = \begin{bmatrix} -e, & \mathbf{P}^T \\ -\mathbf{U}, & \boldsymbol{\varphi} \end{bmatrix}$$

where $\mathcal{P}$, e are scalars, $\mathbf{F}^T$, $\mathbf{P}^T$ are linear forms, $\mathbf{U}$ is a vector and $\boldsymbol{\varphi}$ is a 3×3 tensor in space $\mathcal{D}_t$. In components, $i, j \in \{1, 2, 3\}$

$$S_0 = -\mathcal{P}, \quad S_i = F_i, \quad T_0^0 = -e, \quad T_0^i = -U^i, \quad T_i^0 = P_i, \quad T_i^j = \varphi_i^j$$

$\mathbf{S}^T$ and $\mathbf{T}$ will be given by behavior laws as function of the motion.

For the sake of simplicity, we only consider a first gradient model[5]. However, a second gradient model based on a functional of the form

$$\delta a = \int_{\mathcal{W}} \left[S_i \zeta^i + T_i^j \zeta_{,j}^i + C_i^{jk} \zeta_{,jk}^i \right] dw$$

might be easy to develop as it has been done in statics and would introduce not only new types of forces but also non-usual kinematic quantities [4].

2.2 Physical dimensions of tensor quantities

Manifolds $\mathcal{W}_0$ and $\mathcal{W}$ being contained in no-metric affine spaces, we specify the physical dimensions of the previously defined quantities

Quantity	Dimension	Physical nature
δa	ML^2T^{-1}	Action
dw	L^3T	Volume element of space-time
$\mathcal{P}$	$ML^2T^{-3}.L^{-3}$	Power per volume
$\mathbf{F}^T$	$MLT^{-2}.L^{-3}$	Volume force
e	$ML^2T^{-2}.L^{-3}$	Volume energy
$\mathbf{U}$	$ML^2T^{-2}.LT^{-1}.L^{-3}$	Volume energy flux
$\mathbf{P}^T$	$MLT^{-1}.L^{-3}$	Volume momentum
$\boldsymbol{\varphi}$	$MLT^{-2}.L^{-2}$	Stress tensor

We give to these quantities the following names

Source $\mathbf{S}^T$	Total energy e
Energy source $\mathcal{P}$	Energy flux vector $\mathbf{U}$
Momentum source $\mathbf{F}^T$	Momentum $\mathbf{P}^T$
Stress – energy tensor $\mathbf{T}$	Total stress φ

2.3 The principle of virtual action

Principle 1 *For any virtual displacement, null on the edge $\partial\mathcal{W}$ of $\mathcal{W}$, the virtual action is zero.*

i.e.,

$$\forall \mathbf{z} \in \mathcal{W} \longrightarrow \zeta, \quad \text{with} \quad \zeta = \mathbf{0} \quad \text{if} \quad \mathbf{z} \in \partial\mathcal{W}, \quad \text{then} \quad \delta a = 0$$

This can also be stated in an equivalent way

For any virtual motion with compact support, the virtual action is zero.

3 The equations of dynamics

3.1 Equations of energy and motion

In this section, we consider $\mathbf{T}$ as a differentiable tensor. By integration by parts we obtain

$$\delta a = \int_{\mathcal{W}} [\mathbf{S}^T - \text{Div}(\mathbf{T})] \zeta \, dw + \int_{\partial\mathcal{W}} \mathbf{N}^T \mathbf{T} \zeta \, d\tau \quad (3.3)$$

or

$$\delta a = \int_{\mathcal{W}} [S_i - T_{i,j}^j] \zeta^i \, dw + \int_{\partial\mathcal{W}} N_j T_i^j \zeta^i \, d\tau \quad \{i, j = 0, 1, 2, 3\}$$

Covector $\mathbf{N}^T$ ($N_j \{j = 0, 1, 2, 3\}$) is the annulator of vectors belonging to the hyperplane tangent to $\partial\mathcal{W}$, edge of $\mathcal{W}$ with *surface element* $d\tau$.

The fundamental lemma of calculus of variations associated with the principle of virtual action gives the equation

$$\text{Div} \mathbf{T} = \mathbf{S}^T \quad \text{or} \quad T_{i,j}^j = S_i \quad \{i, j = 0, 1, 2, 3\}, \quad (3.4)$$

where Div denotes the divergence operator in space–time $\mathcal{W}$. This equation can be separated in terms of time and space

The term in time

$$\frac{\partial e}{\partial t} + \text{div} \mathbf{U} = \mathcal{P} \quad \text{or} \quad \frac{\partial e}{\partial t} + U_{,i}^i = \mathcal{P} \quad \{i = 1, 2, 3\},$$

where div denotes the divergence operator in space $\mathcal{D}_t$, is the equation of energy. The energy of a continuous medium is described by a quadri-vector $\mathbf{E} = \begin{bmatrix} e \\ \mathbf{U} \end{bmatrix}$, whose divergence in space-time is the source of energy $\text{Div } \mathbf{E} = \mathcal{P}$ (energy introduced per unit of time into unit of volume).
The terme in space is

$$\frac{\partial \mathbf{P}^T}{\partial t} + \text{div } \boldsymbol{\varphi} = \mathbf{F}^T \quad \text{or} \quad \frac{\partial P_i}{\partial t} + \varphi_{i,j}^j = F_i \quad \{i = 1, 2, 3\}$$

which is the momentum equation. The momentum is described by a (3×4) -tensor $\boldsymbol{\pi} = \begin{bmatrix} \mathbf{P}^T \\ \boldsymbol{\varphi} \end{bmatrix}$ whose divergence in space-time $\text{Div } \boldsymbol{\pi} = \mathbf{F}^T$ is the source of momentum (momentum communicated per unit of time into unit of volume). Equations (3.4) are conservation laws in continuum mechanics [10, 11, 12].

3.2 Relation between equation of energy and equation of motion

Application $\mathcal{F}$ in Eq. (1.1) defines a λ -parameterized motion. If its variation only consists in a change of parameter λ , the motion remains unchanged and the corresponding virtual action is null. For such an application $\boldsymbol{\zeta}$ is proportional to space-time velocity $\mathcal{V} = \begin{bmatrix} 1 \\ \mathbf{v} \end{bmatrix}$ (where $\mathbf{v}$ is the spatial velocity) and assumed null on ∂W . The functional (2.2) is identically zero and

$$\left[\mathbf{S}^T - \text{Div } \mathbf{T} \right] \mathcal{V} \equiv 0 \quad (3.5)$$

is an identity which can be written as

$$\frac{\partial e}{\partial t} + \text{div } \mathbf{U} - \mathcal{P} - \left[\frac{\partial \mathbf{P}^T}{\partial t} + \text{div } \boldsymbol{\varphi} - \mathbf{F}^T \right] \mathbf{v} \equiv 0$$

or

$$\frac{\partial e}{\partial t} + U_{,i}^i - \mathcal{P} - \left[\frac{\partial P_i}{\partial t} + \varphi_{i,j}^j - F_i \right] v^i \equiv 0$$

Theorem 2 *The equation of energy is obtained by multiplying the equation of momentum by the velocity $\mathbf{v}$.*

In Section 6, we will see that Eq. (3.5) represents the equation of entropy.

4 Conservative media

For perfect fluids, hyperelastic media in small or large deformation, and more generally for conservative media because the energy is not dissipated by viscosity or thermal conduction, the virtual action can be constructed as follows

Axiom 3 *To any motion $\mathbf{Z} \longrightarrow \mathbf{z} = \mathcal{F}(\mathbf{Z})$ corresponds an action a localized in space-time $\mathcal{W}$ by a density $\mathcal{L}$ per unit volume and time, defined at any point of $\mathcal{W}$ and written in integral form, i.e.,*

$$a = \int_{\mathcal{W}} \mathcal{L} dw$$

The virtual action δa is the variation of action a .

Axiom 4 *The value of $\mathcal{L}$ at point $\mathbf{z}$ is a given function of $\mathbf{z}$, $\mathcal{B} = \frac{\partial \mathbf{z}}{\partial \mathbf{Z}}$ and $\mathbf{Z}$,*

$$\mathcal{L} = L(\mathbf{z}, \mathcal{B}, \mathbf{Z}) \quad (4.6)$$

The invariance of $\mathcal{L}$ with respect to the parameterization implies that $\mathcal{L}$ depends on $\mathbf{Z}$ only through $\mathbf{X}$ and on $\mathcal{B}$ only through $\mathcal{A}$, where $\mathcal{A} = \begin{bmatrix} 1, & \mathbf{0}^T \\ \mathbf{v}, & \mathbf{F} \end{bmatrix}$, i.e,

$$\mathcal{L} = L_0(\mathbf{z}, \mathcal{A}, \mathbf{Z}) \equiv L_1(t, \mathbf{x}, \mathbf{v}, \mathbf{F}, \mathbf{X}) \quad (4.7)$$

Thanks to axioms, $\mathcal{L}$ is a Lagrangian and the equations of motion (3.4) are equations of Lagrange.

The variation of action a yields a source $\mathbf{S}^T$ and a stress-energy tensor $\mathbf{T}$. By taking $\mathcal{L}$ in the form (4.6), we obtain

$$\mathbf{S}^T = \frac{\partial L}{\partial \mathbf{z}}, \quad \mathbf{T} = L + \mathcal{B} \frac{\partial L}{\partial \mathcal{B}} \quad \text{or} \quad S_i = \frac{\partial L}{\partial z_i}, \quad \mathbf{T}_i^j = L \delta_i^j + \mathcal{B}_k^j \frac{\partial L}{\partial \mathcal{B}_k^i}, \quad \{i, j = 0, 1, 2, 3\}$$

where δ_i^j is the Kronecker symbol. If $\mathcal{L}$ is restricted to form (4.7), we have

$$\mathbf{S}^T = \frac{\partial L}{\partial \mathbf{z}}, \quad \mathbf{T} = L + \mathcal{A} \frac{\partial L}{\partial \mathcal{A}} \Pi \quad (4.8)$$

where Π is the projector parallel to $\mathcal{V}$ of space-time $\mathcal{W}$ into space $\mathcal{D}_t$.

$$\Pi = \begin{bmatrix} 0, & \mathbf{0}^T \\ -\mathbf{v}, & \mathbf{1} \end{bmatrix} \quad \text{where} \quad \Pi_0^0 = 0, \quad \Pi_0^j = -v^j, \quad \Pi_i^0 = 0, \quad \Pi_i^j = \delta_i^j,$$

The formulas (4.8) are explained in terms of time and space. If $\mathcal{L}$ is in form L_1 of Eq. (4.7),

$$\mathbf{S}^T = [-\mathcal{P}, \mathbf{F}^T], \quad \text{where} \quad \mathcal{P} = -\frac{\partial L_1}{\partial t}, \quad \mathbf{F}^T = -\frac{\partial L_1}{\partial \mathbf{x}} \quad (4.9)$$

$$\mathbf{T} = \begin{bmatrix} -e, & \mathbf{P}^T \\ -U, & \mathbf{v}\mathbf{P}^T - \boldsymbol{\sigma} \end{bmatrix} \quad (4.10)$$

where $\mathbf{P}^T = \frac{\partial L_1}{\partial \mathbf{v}}$, $e = \mathbf{P}^T \mathbf{v} - L_1$, $\boldsymbol{\sigma} = -L_1 \mathbf{1} - \mathbf{F} \frac{\partial L_1}{\partial \mathbf{F}}$, $U = e \mathbf{v} - \boldsymbol{\sigma} \mathbf{v}$

In stress–energy tensor $\mathbf{T}$, we have replaced $\boldsymbol{\varphi}$ by $\mathbf{v}\mathbf{P}^T - \boldsymbol{\sigma}$. The static stress tensor is $\boldsymbol{\sigma}$ expressed as a function of L_1 , and the Lagrangian $L_1(t, \mathbf{x}, \mathbf{0}, \mathbf{F}, \mathbf{X})$ provides the virtual work expression.

Theorem 5 *For a conservative motion, the velocity quadri-vector $\mathcal{V}$ is the eigenvector of the stress–energy tensor. The eigenvalue is the Lagrangian.*

Since $\mathcal{V} = \mathbf{0}$, the property comes from the second relation (4.8) and the action is independent of the parameter λ . This property of the stress–energy tensor is expressed as energy flux vector $\mathbf{U}$ is $e\mathbf{v} - \boldsymbol{\sigma}\mathbf{v}$, but is not valid when we have heat conduction.

5 The laws of dynamic behavior

The functional we have chosen as virtual action (2.2) allows to introduce a number of notions, specifying their tensorial nature. By (3.4) and (3.5), the principle of virtual action indicates the relations that necessarily link these notions. It is important to note this construction is only a set of conventions which obviously cannot claim to govern the nature. The study of laws of dynamic behavior can only be carried out when appropriate experiments have been done. In fact, for each concerned medium, the laws express the different dynamic quantities we have introduced (in fact $\mathbf{S}^T$ and $\mathbf{T}$) as function of its motion. There is nothing to restrict the diversity of these laws, they only must be compatible with relations (3.4) and (3.5).

Note that the virtual work in statics is deduced from the virtual action when the motion is reduced to rest. The dynamic behavior law then provides the static behavior law.

The case of conservative media is particularly simple since it is sufficient to express a behavior law by giving the Lagrangian. We first examine this case.

5.1 Choice of the Lagrangian

The Lagrangian summarizes the mechanical properties of the continuous medium and also of the space–time.

The Lagrangian depends (see (4.7))

- On t , $\mathbf{x}$ and $\mathbf{v}$ which are quantities defined in space–time,
- On $\mathbf{X}$ which designates the particle position in the reference space $\mathcal{D}_0$
- On $\mathbf{F}$ which is an application from the tangent vector space $T_{\mathbf{X}}(\mathcal{D}_0)$ at $\mathbf{X}$ into the tangent vector space $T_{\mathbf{x}}(\mathcal{D}_t)$ at $\mathbf{x}$.

The value of the Lagrangian varies from one particle to another one, or by changing the axis in $T_{\mathbf{X}}(\mathcal{D}_0)$ and the Lagrangian $\mathcal{L}$ depends on $\mathbf{X}$ and $\mathbf{F}$. If the Lagrangian is independent of $\mathbf{X}$, the medium is homogeneous. If the space-time possesses homogeneity and symmetry, Lagrangian $\mathcal{L}$ does not vary with t and $\mathbf{x}$ or by changing the axes of $T_{\mathbf{x}}(\mathcal{D}_t)$ (which has the effect of changing the components of $\mathbf{v}$ and $\mathbf{F}$).

These properties lead to different models of Lagrangian and experiments conclude which model is the most appropriate. The simplest way is to consider the Lagrangian as invariant by displacement in space and by translation in time. To express this property (called the principle of material indifference), we have to introduce a metric in the space. Then, the Lagrangian is independent of $\mathbf{z}$ and depends only on $\mathbf{v}$ by its scalar square $|\mathbf{v}|^2 = \mathbf{v}^T \mathbf{v}$ and on $\mathbf{F}$ by the Cauchy tensor $\mathbf{G} = \mathbf{F}^T \mathbf{F}$, i.e.,

$$\mathcal{L} = L(|\mathbf{v}|^2, \mathbf{G}, \mathbf{X}) \quad (5.11)$$

In particular, $\mathbf{S}^T = \mathbf{0}$ (no source) and $\boldsymbol{\sigma} = \boldsymbol{\sigma}^T$ (the constraint is symmetric; noting that the notion of symmetry is meaningless as long as there is no metric).

5.2 Interior and exterior of a medium

The study of motions in a terrestrial laboratory shows that the vertical is a privileged direction, which contradicts the previous invariance. In general, the influence of material bodies outside the medium destroys the invariance. This influence is expressed by geometrical conditions imposed at the boundary of the medium or by contact actions whose virtual forms of action are given, and by forces exerted inside the medium. These actions at distance, of which only two types are currently known - gravitational forces and electro-magnetic forces - are introduced in terms of sources whose virtual actions are added to the Lagrangian.

It is known that in general relativity, the influence of the exterior is taken into account by the metric of the space-time which is no longer an affine space and the definitions as well as the general equations remain valid when covariant derivatives are used [9].

6 Equations of energy and motion for continuous media

6.1 Isolated conservative media

We consider a Lagrangian invariant by time translation and space displacement (material indifference) in form (5.11). More precisely, $\mathcal{L}$ is the difference between the kinetic and potential volumetric energies

$$\mathcal{L} = \rho \left(\frac{1}{2} |\mathbf{v}|^2 - \alpha(\mathbf{G}, s) \right) \quad (6.12)$$

With such a Lagrangian, where ρ denotes the density, α the specific energy and s the specific entropy, the principle of virtual action merges with Hamilton's principle [6, 7].

For the first time, deduced from the concept of kinetic energy which introduces an additional variable ρ corresponding to the density, the mass intervenes. We add the mass behavior: the mass of any material portion of the medium is conserved along the motions and is expressed by the equation

$$\text{Div}(\rho \mathcal{V}) = 0$$

In some specialized theories (diffusion, aerosols, porous media, etc.) we have to write there is no conservation of mass,

$$\text{Div}(\rho \mathcal{V}) = k,$$

where k denotes the source of mass.

The potential energy, called internal energy $\rho \alpha(\mathbf{G}, s)$, is chosen in static and is given by the thermodynamics which introduces the specific entropy s which is a new state variable. The entropy s is not a mechanical variable and its variation δs is zero [1]. This fact summarizes the thermodynamics, and the Kelvin temperature is $\theta = \partial \alpha / \partial s$. The additional variable s being introduced, it seems that an equation is missing. This is not the case, Eq. (3.5) will give the evolution of s .

Equations (4.9) and (4.10) yield $\mathbf{S}^T$ and $\mathbf{T}$ issued from Lagrangian (6.12)

$$\left\{ \begin{array}{l} \text{Source : } \mathbf{S} = \mathbf{0} \\ \text{Momentum source : } \mathbf{P} = \rho \mathbf{v} \\ \text{Stress tensor : } \boldsymbol{\sigma} = 2\rho \mathbf{F} \frac{\partial \alpha}{\partial \mathbf{G}} \mathbf{F}^T \\ \text{Total energy : } e = \rho \left(\frac{1}{2} |\mathbf{v}|^2 + \alpha(\mathbf{G}, s) \right) \\ \text{Energy flux vector : } \mathbf{U} = e \mathbf{v} - \boldsymbol{\sigma} \mathbf{v} \end{array} \right.$$

For the special case of compressible fluid $\alpha = \alpha(\rho, s)$ and $\boldsymbol{\sigma} = -p \mathbf{1}$ where $p = \rho^2 \frac{\partial \alpha}{\partial \rho}$ is the thermodynamic pressure.

In the general case, the equations of motion (3.4) and (3.5) reduce to

$$\text{Div } \mathbf{T} = \mathbf{0} \quad \text{and} \quad (\text{Div } \mathbf{T}) \mathcal{V} = 0 \quad (6.13)$$

Let's clarify (6.13) by decomposing $\mathbf{T}$ into $\mathbf{T}_1$ and $\mathbf{T}_2$ coming from kinetic energy and potential energy, respectively

$$\mathbf{T}_1 = \begin{bmatrix} -\frac{1}{2} \rho |\mathbf{v}|^2, & \rho \mathbf{v}^T \\ -\frac{1}{2} |\mathbf{v}|^2 \rho \mathbf{v}, & \rho \mathbf{v} \mathbf{v}^T \end{bmatrix} \quad \mathbf{T}_2 = \begin{bmatrix} -\rho \alpha, & \mathbf{0}^T \\ -\rho \alpha \mathbf{v} + \boldsymbol{\sigma} \mathbf{v}, & -\boldsymbol{\sigma} \end{bmatrix}$$

Then, $\mathbf{a} = \frac{d\mathbf{v}}{dt}$ denoting the acceleration vector,

$$(\text{Div } \mathbf{T}_1)^T = \begin{bmatrix} -\frac{1}{2} |\mathbf{v}|^2 \text{Div}(\rho \mathcal{V}) - \rho \mathbf{v}^T \mathbf{a} \\ \mathbf{v} \text{Div}(\rho \mathcal{V}) + \rho \mathbf{a} \end{bmatrix} \implies (\text{Div } \mathbf{T}_1)^T \mathcal{V} = \frac{1}{2} |\mathbf{v}|^2 \text{Div}(\rho \mathcal{V})$$

$$(\text{Div } \mathbf{T}_2)^T = \begin{bmatrix} -\alpha \text{Div}(\rho \mathcal{V}) - \rho \frac{d\alpha}{dt} + \text{div}(\boldsymbol{\sigma} \mathbf{v}) \\ -(\text{div}(\boldsymbol{\sigma}))^T \end{bmatrix} \implies$$

$$(\text{Div } \mathbf{T}_2)^T \mathcal{V} = -\alpha \text{Div}(\rho \mathcal{V}) - \rho \frac{d\alpha}{dt} + \text{div}(\boldsymbol{\sigma} \mathbf{v}) - \text{div}(\boldsymbol{\sigma}) \mathbf{v} \equiv -\alpha \text{Div}(\rho \mathcal{V}) - \rho \frac{ds}{dt}$$

Hence,

$$\begin{cases} \rho \frac{d\alpha}{dt} + \rho \mathbf{v}^T \mathbf{a} - \text{div}(\boldsymbol{\sigma} \mathbf{v}) + \frac{1}{2} |\mathbf{v}|^2 \text{Div}(\rho \mathcal{V}) = 0 \\ \rho \mathbf{a} = (\text{div } \boldsymbol{\sigma})^T - \text{Div}(\rho \mathcal{V}) \mathbf{v} \end{cases} \quad (6.14)$$

Equation (6.14)¹ is the equation of energy and (6.14)² is the equation of momentum. Here, the equation (3.5) consequence of Eqs. (6.14) yields

$$\rho \theta \frac{ds}{dt} = \left(\frac{1}{2} |\mathbf{v}|^2 - \alpha \right) \text{Div}(\rho \mathcal{V})$$

which is the equation verified by the specific entropy. We have written the equations with terms of mass source and it is necessary to specify the mechanism by which the mass is introduced.

If there is no source of mass, $\text{Div}(\rho \mathcal{V}) = 0$ and equations (6.14) reduce to

$$\rho \frac{d\alpha}{dt} + \rho \mathbf{v}^T \mathbf{a} - \text{div}(\boldsymbol{\sigma} \mathbf{v}) = 0, \quad \rho \mathbf{a} = (\text{div } \boldsymbol{\sigma})^T, \quad \frac{ds}{dt} = 0 \quad (6.15)$$

6.2 Non-isolated conservative medium

Even if this is often neglected, the continuous medium is subject to the influence of exterior and an external action must be added. The contact forces can be taken into account by a virtual action integrated on the edge $\partial \mathcal{W}$ and have an effect on the boundary conditions.

If remote actions only occur through source terms $\mathbf{S}^T = [-\mathcal{P}, \mathbf{F}^T]$ which have to be known, equations (6.15) are replaced by

$$\begin{cases} \rho \frac{d\alpha}{dt} + \rho \mathbf{v}^T \mathbf{a} - \text{div}(\boldsymbol{\sigma} \mathbf{v}) = \mathcal{P} \\ \rho \mathbf{a} = (\text{div } \boldsymbol{\sigma})^T + \mathbf{F} \\ \rho \theta \frac{ds}{dt} = \mathcal{P} - \mathbf{F}^T \mathbf{v} \end{cases} \quad (6.16)$$

In general, the motion is not adiabatic and $(\mathcal{P} - \mathbf{F}^T \mathbf{v})/(\rho\theta)$ is the source of entropy.

An usual case corresponds to external forces deriving from a potential. In this case, it is better to return to Hamilton's principle with a new Lagrangian with the potential $\Omega(\mathbf{z})$ per unit of mass

$$\mathcal{L} = \rho \left(\frac{1}{2} \mathbf{v}^2 - \alpha(\mathbf{G}, s) - \Omega(\mathbf{z}) \right)$$

The Lagrangian is not invariant by spatial translation. Nevertheless, the continuous medium remains conservative. Term $-\rho\Omega$ added to the Lagrangian introduces a source term $\mathbf{S}_3$ and a tensor $\mathbf{T}_3$

$$\mathbf{S}_3 = -\rho \frac{\partial \Omega}{\partial \mathbf{z}} \quad \text{and} \quad \mathbf{T}_3 = \begin{bmatrix} -\rho \Omega & , & \mathbf{0}^T \\ -\rho \Omega \mathbf{v} & , & \mathbf{0} \end{bmatrix}$$

with

$$\mathcal{P} = \rho \frac{\partial \Omega}{\partial t} \quad \text{and} \quad \mathbf{F} = -\rho \text{grad } \Omega$$

We must add to the first member of Eq. (3.4) the term $\text{Div} \mathbf{T}_3 - \mathbf{S}_3^T$. As $(\text{Div} \mathbf{T}_3 - \mathbf{S}_3^T) \mathcal{V} = 0$, equation (3.5) is unchanged and Eqs. (6.16)² and (6.16)³ write

$$\begin{cases} \rho \mathbf{a} = (\text{div } \boldsymbol{\sigma})^T - \rho \text{grad } \Omega \\ \rho \theta \frac{ds}{dt} = 0 \end{cases}$$

If we write the total energy $e = 1/2 \rho |\mathbf{v}|^2 + \rho \alpha + \rho \Omega$, the equation of energy (6.16)¹ can be written

$$\frac{\partial e}{\partial t} + \text{div } \mathbf{U} = \rho \frac{\partial \Omega}{\partial t}, \quad \text{where} \quad \mathbf{U} = e \mathbf{v} - \boldsymbol{\sigma} \mathbf{v}$$

The motion remains adiabatic. This result is consistent with the thermodynamics: the entropy of a gas in a tank does not change if the tank moves in the gravity field.

It is possible to comment on the above results as follows

- Or to obtain the total energy e , one considers the internal energy and an external potential energy which is added to the kinetic energy. Consequently, we have an energy source $\mathcal{P} = \rho \partial \Omega / \partial t$, a momentum source $\mathbf{F} = -\rho \text{grad } \Omega$ and an additional energy flux vector $\rho \Omega \mathbf{v}$.

- Or we only consider the energy in the form $e = 1/2 \rho |\mathbf{v}|^2 + \rho \alpha$ and we have an energy source $\mathcal{P} = \mathbf{F}^T \mathbf{v}$ and a momentum source $\mathbf{F} = -\rho \text{grad } \Omega$. In both cases the entropy source, different from the energy source, is zero.

6.3 Introduction of dissipative terms

A supplementary tensor $\mathbf{T}_4$ added to the stress–energy tensor allows to introduce dissipative effects (viscosity, thermal conduction).

- In the case of viscous motion, a viscosity stress tensor $\boldsymbol{\sigma}_1$ is added to the stress tensor $\boldsymbol{\sigma}$.

- In the case of thermal conduction, a vector $\mathbf{U}_1$ is added to the energy flux $\mathbf{U}$ [13,14].

The stress–energy tensor becomes

$$\mathbf{T} = \begin{bmatrix} -e & \rho \mathbf{v}^T \\ -e \mathbf{v} + (\boldsymbol{\sigma} + \boldsymbol{\sigma}_1) \mathbf{v} - \mathbf{U}_1, & \rho \mathbf{v} \mathbf{v}^T - \boldsymbol{\sigma} - \boldsymbol{\sigma}_1 \end{bmatrix}$$

with $\mathbf{T}_4 = \begin{bmatrix} 0 & \mathbf{0}^T \\ \boldsymbol{\sigma}_1 \mathbf{v} - \mathbf{U}_1, & -\boldsymbol{\sigma}_1 \end{bmatrix}$.

If $\mathbf{U}_1$ is zero, note that $\mathcal{V}$ remains an eigenvector of the stress–energy tensor. We deduce

$$(\text{Div } \mathbf{T}_4) \mathcal{V} = \Psi - \text{div } \mathbf{U}_1 \quad \text{with} \quad \Psi = \text{div}(\boldsymbol{\sigma}_1 \mathbf{v}) - \text{div}(\boldsymbol{\sigma}_1) \mathbf{v} \equiv \text{Tr} \left(\boldsymbol{\sigma}_1 \frac{\partial \mathbf{v}}{\partial \mathbf{x}} \right)$$

Scalar Ψ is the dissipation function or the power dissipated by the viscosity per unit volume. In the case of isolated media, the equations of motion are

$$\begin{cases} \frac{\partial e}{\partial t} + \text{div}(e \mathbf{v} - \boldsymbol{\sigma} \mathbf{v}) = \text{div}(\boldsymbol{\sigma}_1 \mathbf{v}) - \text{div } \mathbf{U}_1 \\ \rho \mathbf{a} = \text{div}(\boldsymbol{\sigma}^T + \boldsymbol{\sigma}_1^T) \\ \rho \theta \frac{ds}{dt} = \Psi - \text{div } \mathbf{U}_1 \end{cases} \quad (6.17)$$

to which it may be necessary to add source terms for non-isolated media.

7 Behavior laws for dissipative forces

The equations (6.17) only constitute a general framework. It is necessary to present behavior laws characterizing the medium and allowing to obtain $\boldsymbol{\sigma}_1$ and $\mathbf{U}_1$ as function of the motion. The second law of thermodynamics is the only way to limit the possible variety of the behaviors [15].

The entropy of isolated systems can only increase (or remain constant). Another restriction is that the entropy increase occurs at each point, i.e.,

$$\forall \mathbf{z} \in \mathcal{W}, \quad \frac{ds}{dt} \geq 0$$

If we also require that each term constituting the entropy variation remains positive at each point and at each time, we obtain

$$\Psi \geq 0 \quad \text{and} \quad \operatorname{div} \mathbf{U}_1 \leq 0$$

It remains to classify dissipative media, according to the behavior they may have with respect to spatial and material invariance. We only describe the simplest of dissipative media, i.e., the viscous fluids.

1°) $\boldsymbol{\sigma}_1$ is an isotropic function of the deformation rate $\mathbf{D}$ so that

$$\mathbf{D} = \frac{1}{2} \left(\frac{\partial \mathbf{v}}{\partial x} + \left(\frac{\partial \mathbf{v}}{\partial x} \right)^T \right)$$

In the linearised case,

$$\boldsymbol{\sigma} = \lambda \operatorname{Tr}(\mathbf{D}) + 2\mu \mathbf{D},$$

and the dissipation function is

$$\Psi = \lambda (\operatorname{Tr} \mathbf{D})^2 + 2\mu \operatorname{Tr}(\mathbf{D}^2)$$

The condition $\Psi \geq 0$ reduces to $3\lambda + 2\mu \geq 0$ and $\mu \geq 0$.

Terms λ and μ are the dynamic viscosity coefficients, whose values can be found for the different gases and liquids in the tables of physical constants. For the sake of simplicity, the Stokes hypothesis $3\lambda + 2\mu = 0$ is sometimes used, which assumes that there is no energy dissipation if a gas is expanding spherically.

For incompressible fluids, $\operatorname{Tr}(\mathbf{D}) = 0$; then $\boldsymbol{\sigma}_1 = 2\mu \mathbf{D}$, $\Psi = 2\mu \operatorname{Tr}(\mathbf{D}^2)$ and only the μ coefficient is involved.

2°) The vector $\mathbf{U}_1$ represents the energy flux coming from a non-mechanical but thermal action. Since the heat flux goes from hot to cold, for reasons of isotropy we take $\mathbf{U}_1$ proportional to the temperature gradient

$$\mathbf{U}_1 = -k \operatorname{grad} \theta,$$

where $k > 0$ is the heat conduction coefficient. The entropy equation writes:

$$\rho \theta \frac{ds}{dt} = \Psi + \operatorname{div} (k \operatorname{grad} \theta)$$

8 Shock waves

For conservative media, the principle of virtual action is reduced to Hamilton's principle which is valid for the case of shocks [6, 7]. A shock wave is a moving surface, thus a three-dimensional manifold Σ in the space-time $\mathcal{W}$, on which the stress-energy tensor is discontinuous. We can always write the principle of virtual action. The integration by parts expressed in Section 3 relation (3.3) must be modified as follows

$$\delta a = \int_{\mathcal{W}} [\mathbf{S}^T - \operatorname{Div}(\mathbf{T})] \boldsymbol{\zeta} dw + \int_{\partial \mathcal{W}} \mathbf{N}^T \mathbf{T} \boldsymbol{\zeta} d\tau + \int_{\Sigma} \mathbf{N}^T [\mathbf{T}_2 - \mathbf{T}_1] \boldsymbol{\zeta} d\tau$$

In the integral on Σ , $\mathbf{T}_1$ and $\mathbf{T}_2$ are the values of $\mathbf{T}$ on each side of Σ , $\mathbf{N}^T$ is the covector, annulator of vectors in the hyperplane tangent to Σ .

We have $\mathbf{N}^T = [-D_n, \mathbf{n}^T]$, where $\mathbf{n}$ is the downstream unit normal vector to the shock wave and D_n is normal velocity of the shock wave.

We deduce from the principle of virtual action

1°) Outside from Σ , the equations of motion (3.4) are unchanged.

2°) On Σ we have the shock condition $\mathbf{N}^T [\mathbf{T}_2 - \mathbf{T}_1] = \mathbf{0}^T$.

We note that on Σ we automatically obtain the conservation of momentum and the conservation of energy which is classically postulated [6].

9 Conclusion and comments

The principle of virtual work is convenient for studying the statics of continuous media. The principle of virtual action is suitable for dynamics.

In this paper, we present the classical case of fluid mechanics and elasticity. This presentation can be extended to more complex media as it is done by the principle of virtual work in the case of second gradient models, capillary fluids, metamaterials, etc [16,17,18]. The principle of virtual action allows us to directly obtain the equations of motion and energy and makes explicit the possibility of shock waves for dispersive media [8].

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